Gauge invariance of thermal transport coefficients

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This paper is dedicated to our friend and distinguished colleague Flavio Toigo, a long-time enthusiast of hydrodynamic fluctuations, on the occasion of his seventieth birthday.

November 30, 2015

Abstract Thermal transport coefficients are independent of the specific microscopic expression for the energy density and current from which they can be derived through the Green-Kubo formula. We discuss this independence in terms of a kind of gauge invariance resulting from energy conservation and extensivity, and demonstrate it numerically for a Lennard-Jones fluid, where different forms of the microscopic energy density lead to different time correlation functions for the heat flux, all of them, however, resulting in the same value for the thermal conductivity.

PACS 65.20.De, 66.10.cd, 66.30.Xj, 66.70.-f

Keywords Thermal conductivity, heat transport, hydrodynamic fluctuations, molecular dynamics, Green Kubo

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It has long been thought that the inherent indeterminacy of any quantum mechanical expression for the energy density would hinder the evaluation of thermal transport coefficients from equilibrium ab-initio molecular dynamics (AIMD), using the Green-Kubo (GK) formalism [1,2,3,4]. In classical molecular dynamics (CMD) this goal is achieved by decomposing the total energy of an extended system into localized atomic contributions and by deriving from this decomposition an explicit (and allegedly unique) expression for the energy flux. While the calculation of thermal transport coefficients from equilibrium AIMD has been successfully addressed by some of us in a recent work [5], the question still remains as to whether the expression for the energy flux currently used in CMD is uniquely defined and, in the negative, how is it that different definitions of the energy flux would lead to the same value for the thermal conductivity. In this paper we show that different equivalent definitions for the atomic energies in a classical system lead to different expressions for the macroscopic energy flux, and demonstrate numerically in the case of a Lennard-Jones fluid that these expressions result in the same value for the thermal conductivity, as evaluated from equilibrium CMD through the GK formula. This finding is then rationalized in terms of a kind of gauge invariance of heat transport coefficients, resulting from energy conservation and extensivity.

According to the GK formalism [1,2,3,4], the heat conductivity κ of an isotropic material can be expressed in terms of the auto-correlation function of the macroscopic heat flux, $\mathbf{J}_q(t)$, as:

$$\kappa = \frac{1}{3Vk_BT^2} \int_0^\infty \langle \mathbf{J}_q(t) \cdot \mathbf{J}_q(0) \rangle dt, \tag{1}$$

where brackets $\langle \cdot \rangle$ indicate canonical averages, k_B is the Boltzmann constant, and V and T are the system volume and temperature, respectively. The heat flux is the macroscopic average of the heat current density, which is in turn defined as the non-convective component of the energy current density. Atoms in solids can be assumed to not diffuse, while in one-component and molecular fluids convective energy transport can be disregarded because of momentum conservation. Because of this, in the following we assume that energy and heat currents coincide.

In CMD the macroscopic energy flux is expressed in terms of suitably defined atomic energies whose sum yields the total energy of the system. For the sake of simplicity, we restrict our attention to one-component systems held together by pair potentials, in which case the atomic energies can be defined as [6]:

$$\epsilon_I(\mathbf{R}, \mathbf{V}) = \frac{1}{2} M V_I^2 + \frac{1}{2} \sum_{I \neq I} v(|\mathbf{R}_I - \mathbf{R}_J|), \tag{2}$$

where M is the atomic mass, $\mathbf{R} \doteq \{\mathbf{R}_I\}$ and $\mathbf{V} \doteq \{\mathbf{V}_I\}$ are atomic coordinates and velocities, respectively, v(R) is the inter-atomic pair potential, and the indices I and J run over all the atoms in the system. Using standard manipulations [6], the macroscopic energy flux can be obtained from Eq. (2) as:

$$\mathbf{J}_e = \sum_{I} \epsilon_I \mathbf{V}_I + \frac{1}{2} \sum_{I,J \neq I} (\mathbf{V}_I \cdot \mathbf{F}_{IJ}) (\mathbf{R}_I - \mathbf{R}_J), \tag{3}$$

where $\mathbf{F}_{IJ} = -\nabla v(|\mathbf{R}_I - \mathbf{R}_J|)$ is the force exerted by atom J on atom I. It is often implicitly assumed that the well-definedness of thermal transport coefficients



Fig. 1 The energy of an isolated system is the sum of the energies of its subsystems (as defined when they are isolated as well) plus the interaction among them, W, whose magnitude scales as the area of the interface, depicted in red. When defining the energies of individual subsystems, \mathcal{E} , W has to be arbitrarily partitioned among them.

would stem from the uniqueness of the decomposition of the system's total energy into localized, atomic, contributions. This assumption is manifestly incorrect, as any decomposition leading to the same value for the total energy as Eq. (2) should be considered as legitimate. The difficulty of partitioning a system's energy into subsystems' contributions is illustrated in Fig. 1, which depicts a system made of two interacting subsystems. When defining the energy of each of the two susbsystems, an arbitrary decision has to be made as to how the interaction energy is partitioned. In the case depicted in Fig. 1, for instance, the energy of each of the two subsystems can be defined as $\mathcal{E}(\Omega_i) = E(\Omega_i) + \frac{1}{2}(1 \pm \gamma)W_{12}$, where $E(\Omega_i)$ are the energies of the two isolated subsystems, W_{12} their interaction energy, and γ an arbitrary constant. In the thermodynamic limit, when the energy of any relevant subsystem is much larger than the interaction between any pairs of them, the value of the γ constant is irrelevant. When it comes to defining energy densities (i.e. energies of infinitesimal portions of a system) or atomic energies, instead, the magnitude of the interaction between different susbsystems is comparable to the their energies, which become therefore intrinsically ill-defined.

As a specific example, we consider the following definition for the atomic energies [7]:

$$\epsilon_I^{\Gamma}(\mathbf{R}, \mathbf{V}) = \frac{1}{2} M_I V_I^2 + \frac{1}{2} \sum_{J \neq I} v(|\mathbf{R}_I - \mathbf{R}_J|) (1 + \Gamma_{IJ}), \tag{4}$$

where $\Gamma_{IJ}=-\Gamma_{JI}$ is any antisymmetric matrix. As the inter-atomic potential appearing in Eq. (4) is symmetric with respect to the atomic indices, it is clear that the sum of all the atomic energies does not depend on Γ , thus making any choice of Γ equally permissible. This trivial observation has deep consequences on the theory of thermal fluctuations and transport, because the value of the macroscopic energy flux, instead, depends explicitly on Γ , thus making one fear that the resulting transport coefficients would depend on Γ as well. Using the same manipulations that lead from Eq. (2) to Eq. (3), for any choice of the Γ matrix in Eq. (4), a corresponding expression for the macroscopic energy flux can be found, reading [7]:

$$\mathbf{J}_{e}^{\Gamma} = \mathbf{J}_{e} + \frac{1}{2} \sum_{I,J \neq I} \Gamma_{IJ} \left[v(|\mathbf{R}_{I} - \mathbf{R}_{J}|) \mathbf{V}_{I} - (\mathbf{V}_{I} \cdot \mathbf{F}_{IJ}) (\mathbf{R}_{I} - \mathbf{R}_{J}) \right]. \tag{5}$$

In order to illustrate this state of affairs, we have performed CMD simulations for a fluid of identical atoms, interacting through a Lennard-Jones potential: $v(R) = 4\varepsilon \left[\left(\frac{\sigma}{R}\right)^{12} - \left(\frac{\sigma}{R}\right)^6 \right]$ at density-temperature conditions $\rho = 0.925\sigma^{-3}$ and

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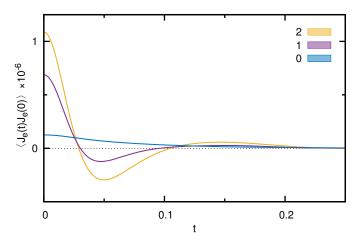


Fig. 2 Time correlation functions of the modified macroscopic energy flux, as defined in Eq. (5), for different definitions of the Γ matrix (see text). The "0" line refers to the original definition (i.e. $\Gamma=0$), whereas the label "1" and "2" correspond to the two definitions of the Γ matrix given in Eq. (6). The parameters used are $\gamma_1=10$ and $\gamma_2=2.5$. Error bars, as estimated by standard block analysis, are smaller than the thickness of the lines. Units are Lennard-Jones units ($M=\sigma=\varepsilon=1$).

 $T=1.86\varepsilon/k_B$, using cubic simulation cells containing 256 atoms in the iso-choric microcanonical ensemble, (NVE) [8]. In Fig. 2 we display the macroscopic energy-flux autocorrelation function corresponding to different choices of the Γ matrix in Eqs. (4) and (5). The Γ matrices have been constructed in two different ways, according to the prescriptions:

$$\Gamma_{IJ} = \begin{cases}
\frac{1}{2} \left(A_{IJ} - A_{JI} \right) & \text{where the matrix elements of } A \text{ are drawn from a uniform deviate in the } \\
[0, \gamma] & \text{interval.}
\end{cases}$$

$$0, +\gamma, -\gamma \quad \text{according to whether } I = J, I > J, \text{ or } \\
I < J.$$
(6)

Fig. 2 clearly shows that the $\langle \mathbf{J}_e^{\Gamma}(t) \cdot \mathbf{J}_e^{\Gamma}(0) \rangle$ correlation functions dramatically depend on the Γ matrices in Eqs. (4) and (5). Notwithstanding, the integrals of all these time correlation functions tend to the same limit at large integration times, as displayed in Fig. 3.

In order to get insight into this remarkable invariance property, let us inspect the difference between the generalized flux in Eq. (5) and the standard expression of Eq. (3):

$$\Delta \mathbf{J}_e^{\Gamma} = \mathbf{J}_e^{\Gamma} - \mathbf{J}_e = \frac{\mathrm{d}}{\mathrm{dt}} \frac{1}{4} \sum_{I,J \neq I} \Gamma_{IJ} v(|\mathbf{R}_I - \mathbf{R}_J|) (\mathbf{R}_I - \mathbf{R}_J). \tag{7}$$

We see that the two different expressions for the macroscopic energy flux differ by a total time derivative. In the following, we will show that this is a consequence of energy conservation and extensivity and a sufficient condition for the corresponding thermal conductivity to coincide.

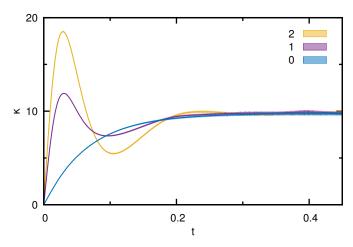


Fig. 3 Integral of the time correlation functions displayed in Fig. 2, multiplied by the prefactor appearing in the GK relation, Eq. (1), as a function of the upper limit of integration. Units are Lennard-Jones units (see caption to Fig. 2). The barely visible shaded area surrounding each line is an indication of the error bars, as estimated by standard block analysis.

The very possibility of defining an energy current density stems from energy extensivity and conservation. Energy is extensive: because of this, the energy of a macroscopic sample of matter of volume Ω can be written as the integral of an energy density, $e(\mathbf{r})$:

$$E[\Omega] = \int_{\Omega} e(\mathbf{r}) d\mathbf{r}.$$
 (8)

Of course, the energy density appearing in Eq. (8) is not uniquely defined, the only requirement being that its integral over a domain Ω is well defined in the thermodynamic limit, *i.e.* two different densities whose integral over a domain Ω differ by a quantity that scales as the area of the domain boundary should be considered as equivalent. This equivalence can be expressed by the condition that two equivalent densities, say $e_1(\mathbf{r})$ and $e_2(\mathbf{r})$, differ by the divergence of a (bounded) vector field:

$$e_2(\mathbf{r}) = e_1(\mathbf{r}) + \nabla \cdot \mathbf{p}(\mathbf{r}).$$
 (9)

In a sense, two equivalent energy densities can be thought of as different gauges of the same scalar field.

Energy is also conserved: because of this, for any given gauge of the energy density, $e(\mathbf{r})$, an energy current density can be defined, $\mathbf{j}_e(\mathbf{r},t)$, so as to satisfy the continuity equation:

$$\dot{e}(\mathbf{r},t) = -\nabla \cdot \mathbf{j}_e(\mathbf{r},t),\tag{10}$$

where the dot indicates a time derivative. By combining Eqs. (9) and (10) we see that energy current densities and macroscopic fluxes transform under a gauge transformation as:

$$\mathbf{j}_2(\mathbf{r},t) = \mathbf{j}_1(\mathbf{r},t) - \dot{\mathbf{p}}(\mathbf{r},t), \tag{11}$$

$$\mathbf{J}_2(t) = \mathbf{J}_1(t) - \dot{\mathbf{P}}(t),\tag{12}$$

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where $\mathbf{P}(t) = \int \mathbf{p}(\mathbf{r}, t) d\mathbf{r}$. We conclude that the macroscopic energy fluxes in two different energy gauges differ by the total time derivative of a vector.

Our previous findings on the energy flux of a system of classical atoms interacting through pair potentials as embodied in Eq. (7) can be recovered by defining the corresponding energy density as:

$$e(\mathbf{r}) = \sum_{I} \delta(\mathbf{r} - \mathbf{R}_{I}) \epsilon_{I}. \tag{13}$$

By taking the first moment of the continuity equation, Eq. (10), with respect to \mathbf{r} and integrating by parts its right-hand side, one sees that the macroscopic average of the energy current density is the first moment of the time derivative of the energy density:

$$\mathbf{J}_{e}(t) = \int \dot{e}(\mathbf{r}, t)\mathbf{r} \, d\mathbf{r}. \tag{14}$$

Eq. (14) is ill-defined in periodic boundary conditions essentially for the same reasons why macroscopic polarization in dielectrics is so [9]. By plugging Eq. (13) into Eq. (14), using Newton's equations of motion, and reducing the resulting expression to a boundary-insensitive form, one easily arrives at the expressions for the macroscopic energy flux given by Eqs. (3) and (5).

We now show that the energy fluxes of the same system in two different energy gauges, e_1 and e_2 , thus differing by a total time derivative, as in Eq. (12), result in the same heat conductivity, as given by the Green-Kubo formula, Eq. (1). Let us indicate by κ_1 and κ_2 the thermal conductivities in the two gauges. Using Eq. (12) and the property that classical time auto-correlation functions are even in time, one obtains:

$$\kappa_2 = \kappa_1 + \frac{1}{6Vk_BT^2} \int_{-\infty}^{+\infty} \frac{d}{dt} \left(\langle \mathbf{P}(-t) \cdot \mathbf{J}_1(0) \rangle - \langle \mathbf{P}(t) \cdot \mathbf{J}_1(0) \rangle + \langle \mathbf{P}(t) \cdot \dot{\mathbf{P}}(0) \rangle \right) dt. \quad (15)$$

The integral on the right-hand side of Eq. (15) vanishes because the correlation function of two observables at large time lags factorizes into the product of two time-independent average values and because the average value of the current J_1 , as well as of any total time derivative, vanishes at equilibrium. We conclude that the heat conductivities computed in different energy gauges coincide, as they must on physical grounds.

In this paper we have demonstrated that, while the heat flux is inherently undetermined at the atomic level, the heat conductivity resulting from it through the Green-Kubo formula is indeed well defined, as any measurable property must be. This indeterminacy stems from the liberty one has to formally unpack the total energy of an extended system into localized contributions in an infinite number of equivalent ways. We believe that this freedom can be exploited to design the definition of the local energy (be it in terms of atomic energies or energy densities), so as to optimise the convergence of computer simulations, regarding simulation length, system size, or both.

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